

Fake Supergravity

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Supersymmetry

- ▶ Space-time symmetry algebra (Poincaré) extended with **fermionic** (anticommuting) generators $Q_{i\alpha}$, $i = 1, \dots, \mathcal{N}$
- ▶ The Q generators transform as spinors under the Poincaré group, and

$$\{Q, Q\} \propto P$$

- ▶ Action on *Bosonic* and *Fermionic* fields:

$$\delta_Q(\epsilon)B \sim \epsilon F, \quad \delta_Q(\epsilon)F \sim \epsilon B$$

- ▶ Every bosonic field must have a fermionic *superpartner* and viceversa

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- ▶ Every bosonic field must have a fermionic *superpartner* and viceversa
- ▶ The supersymmetry algebra may be further extended with internal symmetry generators T_A (**R-symmetry**), with

$$[T, Q] \propto Q$$

Supergravity

- ▶ We could require a theory to be invariant under **local** supersymmetry transformations: $\epsilon \rightarrow \epsilon(x)$
- ▶ Since $\{Q, Q\} \propto P$ the theory must also be invariant under local translations (diffeomorphisms)
- ▶ General relativistic theory: includes gravity \implies **supergravity**

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Example: $\mathcal{N} = 1$ minimal supergravity in 4 dimensions

$$e^{-1}\mathcal{L} = R - \bar{\psi}_\mu \gamma^{\mu\nu\rho} D_\nu \psi_\rho$$
$$\delta e_\mu^a = \frac{1}{2} \bar{\epsilon} \gamma^a \psi_\mu, \quad \delta \psi_\mu = D_\mu \epsilon \equiv \partial_\mu \epsilon + \frac{1}{4} \omega_{\mu ab} \gamma^{ab} \epsilon$$

Solving einstein equations

Einstein equations

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi T_{\mu\nu} - \Lambda g_{\mu\nu}$$

- ▶ 10 second order partial differential equations
- ▶ Highly coupled
- ▶ Must be solved together with matter fields equations

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Finding analytical solutions is not an easy task

Usually solved by first imposing some symmetry on the solution

Supersymmetric solutions

If we consider a supergravity theory, we can look for **supersymmetric solutions**:

- ▶ purely bosonic \rightarrow they are solutions both of the full theory and of the truncated bosonic sector
- ▶ invariant under some (global) supersymmetry transformation

The supersymmetry variation of such field configurations must vanish

- ▶ The variation of the bosonic fields automatically vanishes:
 $\delta_{\mathcal{Q}}(\epsilon)B \sim \epsilon F = 0$
- ▶ For the fermionic fields we must impose $\delta_{\mathcal{Q}}(\epsilon)F = 0$

Killing spinor equations

In any supergravity there is at least one gravitino field

- ▶ Gravitino condition: $\delta_{\mathcal{Q}}(\epsilon)\psi_{\mu}^i = \hat{D}_{\mu}\epsilon^i = 0$
- ▶ Other fermionic fields: $\delta_{\mathcal{Q}}(\epsilon)\zeta^A = 0, \dots$

These conditions depend on the bosonic fields and are called **Killing spinor equations**; the solutions $\epsilon^i(x)$ are **Killing spinors**.

- ▶ First order differential equations \rightarrow easier to solve
- ▶ The integrability conditions impose strong constraints on the field configuration \rightarrow usually they imply the equations of motion

Fake Supergravity

Can this method be generalized to theories that aren't truncations of supersymmetric ones?

e.g.: typically SUGRAs admit only negative or vanishing cosmological constant

Old idea: pre-WWII “Wave Geometry” program in Hiroshima

- ▶ Inspired by Dirac's equation as “square root” of Klein-Gordon equation
- ▶ Attempt to classify all spacetime metrics admitting a nonzero spinor k satisfying

$$(D_\mu + M_\mu)k = 0$$

for all possible choices of M_μ .

Fake $\mathcal{N} = 2$ gauged pure supergravity

$\mathcal{N} = 2$ pure supergravity in 4d with gauged $U(1)$ R-symmetry subgroup

Bosonic lagrangian:
$$e^{-1}\mathcal{L}_{Bos.} = \frac{1}{2}R - \frac{1}{2}F_{\mu\nu}F^{\mu\nu} + 3g^2$$

KSEs:
$$D_{\mu}\epsilon_i = -\frac{g}{2}\gamma_{\mu}\epsilon_{ij}e^j - igA_{\mu}\epsilon_i + iF_{\mu\nu}^{+}\gamma^b\epsilon_{ij}e^j$$

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$\mathcal{N} = 2$ *fake* pure supergravity in 4d with gauged \mathbb{R} -symmetry

$$\text{Bosonic lagrangian:} \quad e^{-1}\mathcal{L}_{Bos.} = \frac{1}{2}R - \frac{1}{2}F_{\mu\nu}F^{\mu\nu} - 3g^2$$

$$\text{fake KSEs:} \quad D_{\mu}\epsilon_i = -i\frac{g}{2}\gamma_{\mu}\varepsilon_{ij}e^j + gA_{\mu}\epsilon_i + iF_{\mu\nu}^+\gamma^b\varepsilon_{ij}e^j$$

The Kastor-Traschen solution

(D. Kastor and J.H. Traschen 1993)

$$ds^2 = -U^{-2}dt^2 + U^2d\mathbf{r}^2, \quad A = U^{-1}dt$$

$$U = \frac{t}{t_0} + \sum_{i=1}^n \frac{Q_i}{|\mathbf{r} - \mathbf{r}_i|}$$

- ▶ Multi-black hole solution
- ▶ Cosmological
- ▶ Time-dependent

Fake $\mathcal{N} = 2$ gauged sugra coupled to vector multiplets

Bosonic lagrangian:

$$e^{-1} \mathcal{L}_{Bos.} = \frac{1}{2} R - g_{i\bar{j}} \mathcal{D}_\mu z^i \mathcal{D}^\mu \bar{z}^{\bar{j}} - V \\ + \Im(\mathcal{N})_{\Lambda\Sigma} F_{\mu\nu}^\Lambda F^{\Sigma\mu\nu} - \Re(\mathcal{N})_{\Lambda\Sigma} F_{\mu\nu}^\Lambda \star F^{\Sigma\mu\nu}$$

fake KSEs:

$$D_\mu \epsilon_I + \frac{i}{2} Q_\mu \epsilon_I + \frac{ig}{2} A_\mu^\Lambda (P_\Lambda + iC_\Lambda) \epsilon_I = -2i \mathcal{L}_\Lambda F_{\mu\nu}^{\Lambda+} \gamma^\nu \epsilon_{IJ} \epsilon^J - \frac{ig}{4} C_\Lambda \mathcal{L}^\Lambda \gamma_\mu \epsilon_{IJ} \epsilon^J \\ i \not{D} z^i \epsilon^I = \bar{f}_\Lambda^i \not{F}^{\Lambda+} \epsilon^{IJ} \epsilon_J + \frac{ig}{2} \bar{f}^{i\Lambda} (P_\Lambda + iC_\Lambda) \epsilon^{IJ} \epsilon_J$$

- ▶ \mathbb{R} -symmetry gauged with connection $C_\Lambda A^\Lambda$
- ▶ The lagrangian can be completely determined by specifying a *prepotential* and the constants C_Λ
- ▶ Fake supersymmetric solutions: general method by [P. Meessen and A. Palomo-Lozano \(2009\)](#)

Einstein-Maxwell-Scalar fake supersymmetric solutions

$$e^{-1}\mathcal{L} = \frac{1}{2} \left\{ R - \frac{n_0 n_1}{8} \partial_\mu \phi \partial^\mu \phi - \frac{n_0}{4} e^{\frac{n_1}{2} \phi} F_{\mu\nu}^0 F^{0\mu\nu} - \frac{n_1}{4} e^{-\frac{n_0}{2} \phi} F_{\mu\nu}^1 F^{1\mu\nu} \right. \\ \left. - \frac{1}{2} \left[\frac{n_0(n_0 - 1)}{t_0^2} e^{-\frac{n_1}{2} \phi} + 2 \frac{n_0 n_1}{t_0 t_1} e^{\frac{n_0 - n_1}{4} \phi} + \frac{n_1(n_1 - 1)}{t_1^2} e^{\frac{n_0}{2} \phi} \right] \right\}$$

$$ds^2 = -U^{-2} dt^2 + U^2 d\mathbf{r}^2, \quad U = H_0^{\frac{n_0}{4}} H_1^{\frac{n_1}{4}} \equiv \left[\frac{t}{t_0} + \mathcal{H}_0 \right]^{\frac{n_0}{4}} \left[\frac{t}{t_1} + \mathcal{H}_1 \right]^{\frac{n_1}{4}}$$

$$\phi = \log \frac{H_0}{H_1}, \quad A^0 = H_0^{-1} dt, \quad A^1 = H_1^{-1} dt$$

- ▶ One abelian vector multiplet
- ▶ No additional gaugings
- ▶ Prepotential: $F = -\frac{i}{4} (X^0)^{\frac{n_0}{2}} (X^1)^{\frac{n_1}{2}}$, $n_0 + n_1 = 4$
- ▶ Generalization of the solutions by [G.W. Gibbons and K.I. Maeda \(2010\)](#) ($t_0 \rightarrow \infty$ or $t_1 \rightarrow \infty$)

Conclusions

- Summary:
- ▶ (Very) brief introduction to supersymmetry and supergravity
 - ▶ Supersymmetric solutions and Killing spinor equations
 - ▶ Generalization: fake supergravity
 - ▶ Example: Kastor-Traschen solution from $\mathcal{N} = 2$ gauged fake sugra
 - ▶ Example: our solution
- Outlook:
- ▶ Study physical properties of our solution
 - ▶ Generalize to other models
 - ▶ Look for multi-black hole solutions in non-flat FLRW cosmologies